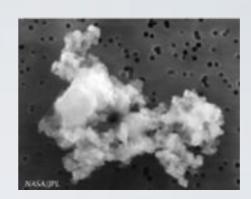
## FLUX OF LARGE METEOROIDS FROM LUNAR IMPACT MONITORING AND INFRASOUND

Bill Cooke Lead, NASA Meteoroid Environments Office william.j.cooke@nasa.gov Meteoroid - small rocky/ icy debris out in space



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Meteor - streak of light/ionization produced by a meteoroid ablating in an atmosphere

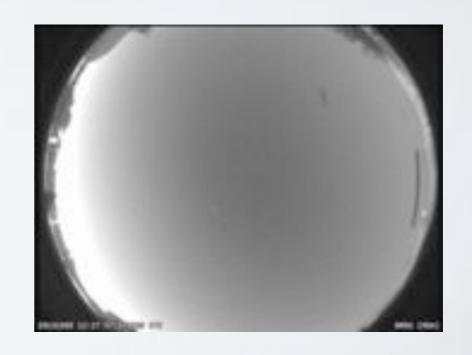
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Meteor - streak of light/ionization produced by a meteoroid ablating in an atmosphere

Fireball - meteor with a peak brightness greater than the planet Venus (m<sub>v</sub> -4). Also called a bolide

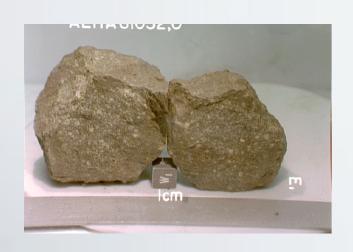




Super bolide - a meteor sufficiently energetic to be detected by seismic/other sensors (m<sub>v</sub> ~-18)



Super bolide - a meteor sufficiently energetic to be detected by seismic/other sensors (m<sub>v</sub> ~-18)



Meteorite - any part of a meteoroid/ asteroid that makes it to the surface

#### ASTEROID OR METEOROID?

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Some scientists say meteoroids are smaller than one meter (beach ball size); others place the limit at 10 meters (garage size)

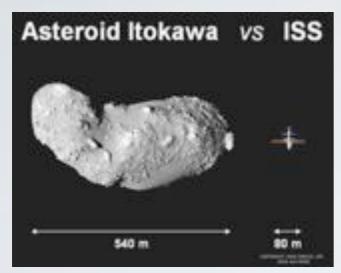
#### ASTEROID OR METEOROID?

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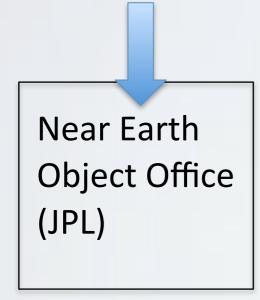
Some scientists say meteoroids are smaller than one meter (beach ball size); others place the limit at 10 meters (garage size)

International Astronomical Union definition: "A meteoroid is a solid object moving in interplanetary space, of a size considerably smaller than an <u>asteroid</u> and considerably larger than an atom."

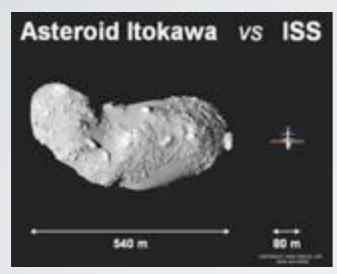
#### How NASA handles meteoroids, asteroids, and space junk



Earth-approaching asteroids



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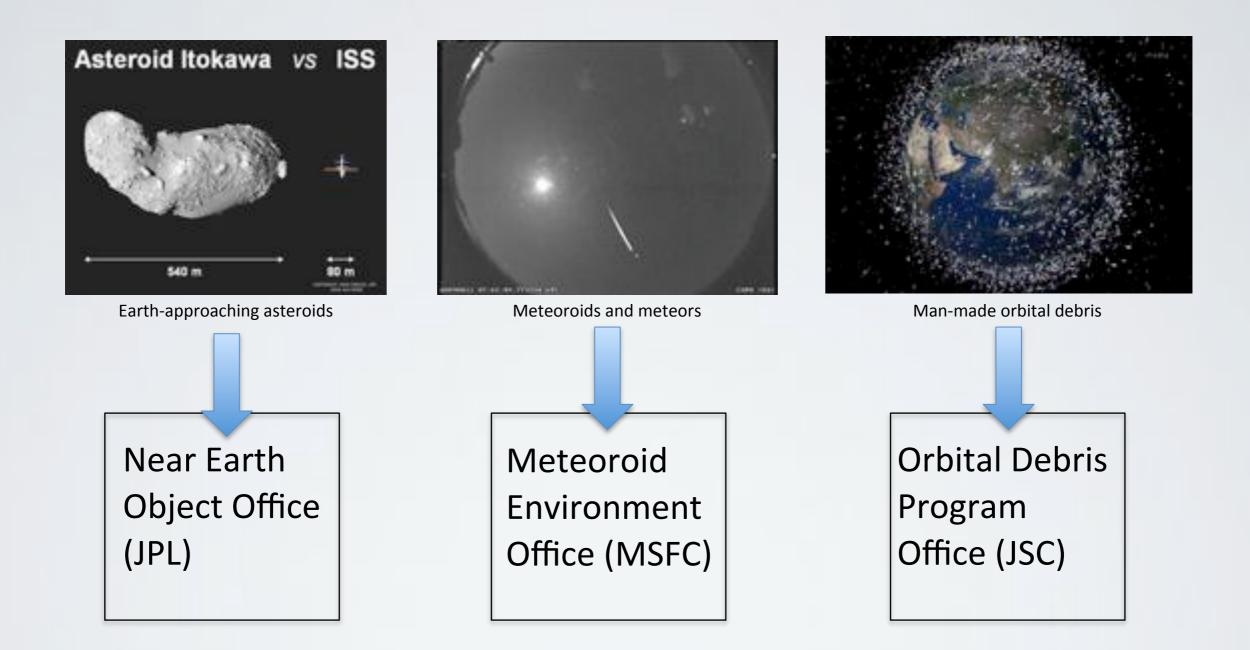
Earth-approaching asteroids





Orbital Debris
Program
Office (JSC)

#### How NASA handles meteoroids, asteroids, and space junk





Develop, maintain, and distribute an sporadic meteoroid model suitable for spacecraft engineering/design

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Provide meteor shower forecasts to NASA/USG spacecraft operators

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Monitor the meteoroid environment in near-Earth space (SSA)

Develop, maintain, and distribute an sporadic meteoroid model suitable for spacecraft engineering/design

Provide meteor shower forecasts to NASA/USG spacecraft operators

Monitor the meteoroid environment in near-Earth space (SSA)

Conduct and manage research to improve sporadic and shower meteoroid models

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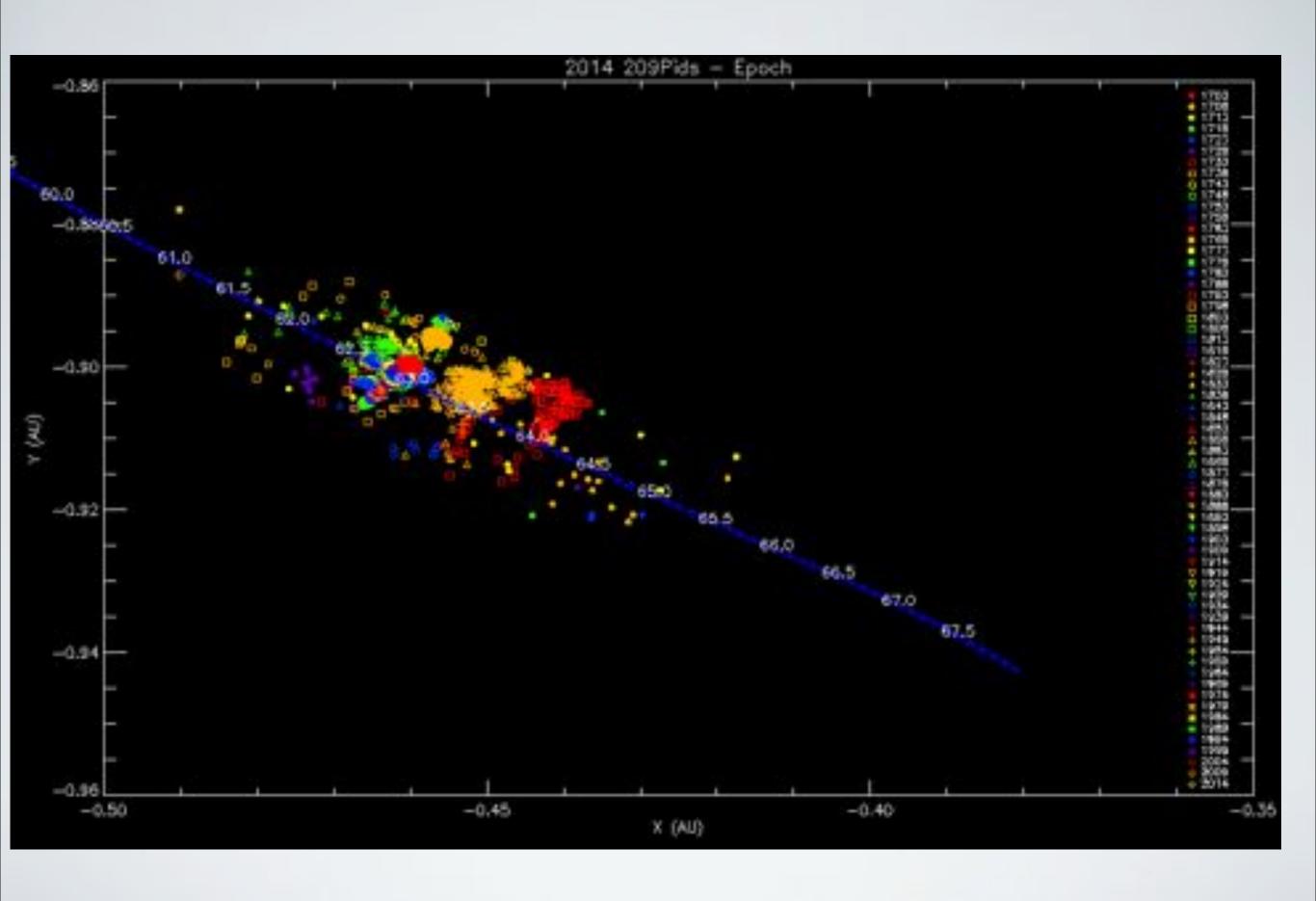
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Models indicate mm size particles in stream, so potential risk to Earth orbiting spacecraft

Model	Time on May 24 (UT)	ZHR (#/hr)	RA, Dec (*)	V <sub>E</sub> (km/s)
Lyytinen & Jenniskens (1929)	3:19			
Lyytinen & Jenniskens (1979)	6:04			
MSFC (peak 1)	6:11			
Lyytinen & Jenniskens (1818, 1853)	6:33			
MSFC (peak 2)	6:56			
Lyytinen & Jenniskens (1903, max)	6:59	15.86	125, +78	15.86
Lyytinen & Jenniskens (1909)	7:15			
Maslov (1898-1919; 1903)	7:18	200-300	122.8, +79.1	16.2
Maslov (max)	7:21	100	122.8, +79.0	
Vaubaillon	7:40	100-400	~125, +79	
Lyytinen & Jenniskens (1914)	7:49			
Maslov (1763-1783)	7:55	50-150	122.8, +79.0	16.2
MSFC (peak 3)	8:10			
Jenniskens (general)			123, +79	16





Monday, March 31, 14

## 7-year observing program

Goal: Monitor the Moon for impact flashes produced by meteoroids striking the lunar surface.



Observation from MSFC

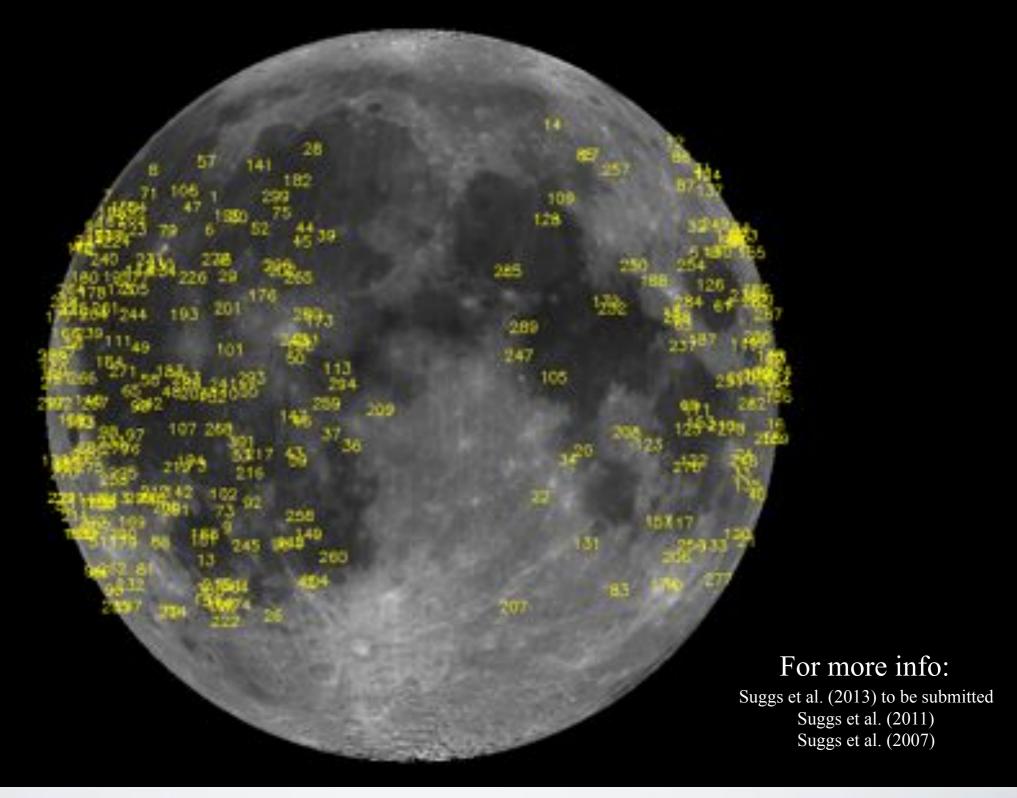
- Two 0.35m telescopes simultaneously
- Black & white CCD video cameras
- Interleaved 30fps video digitized, recorded
- Video analyzed with custom software

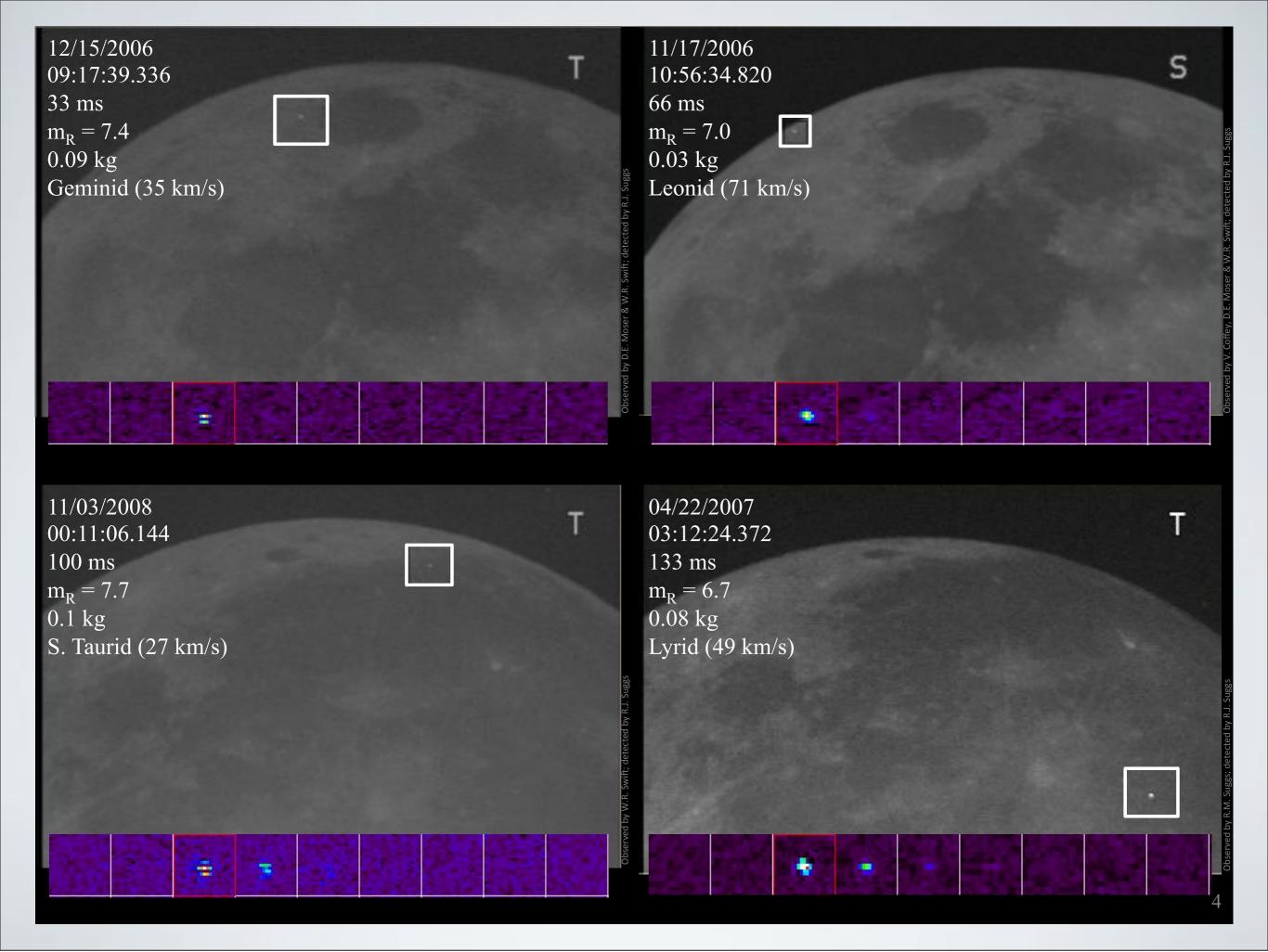


Field of View

- FOV covers approx. 20 arcmin
- $4 \times 10^6$  km<sup>2</sup> on the leading or trailing edge
- Observing when illumination 10-50%
- Maximum 10 observing nights/month

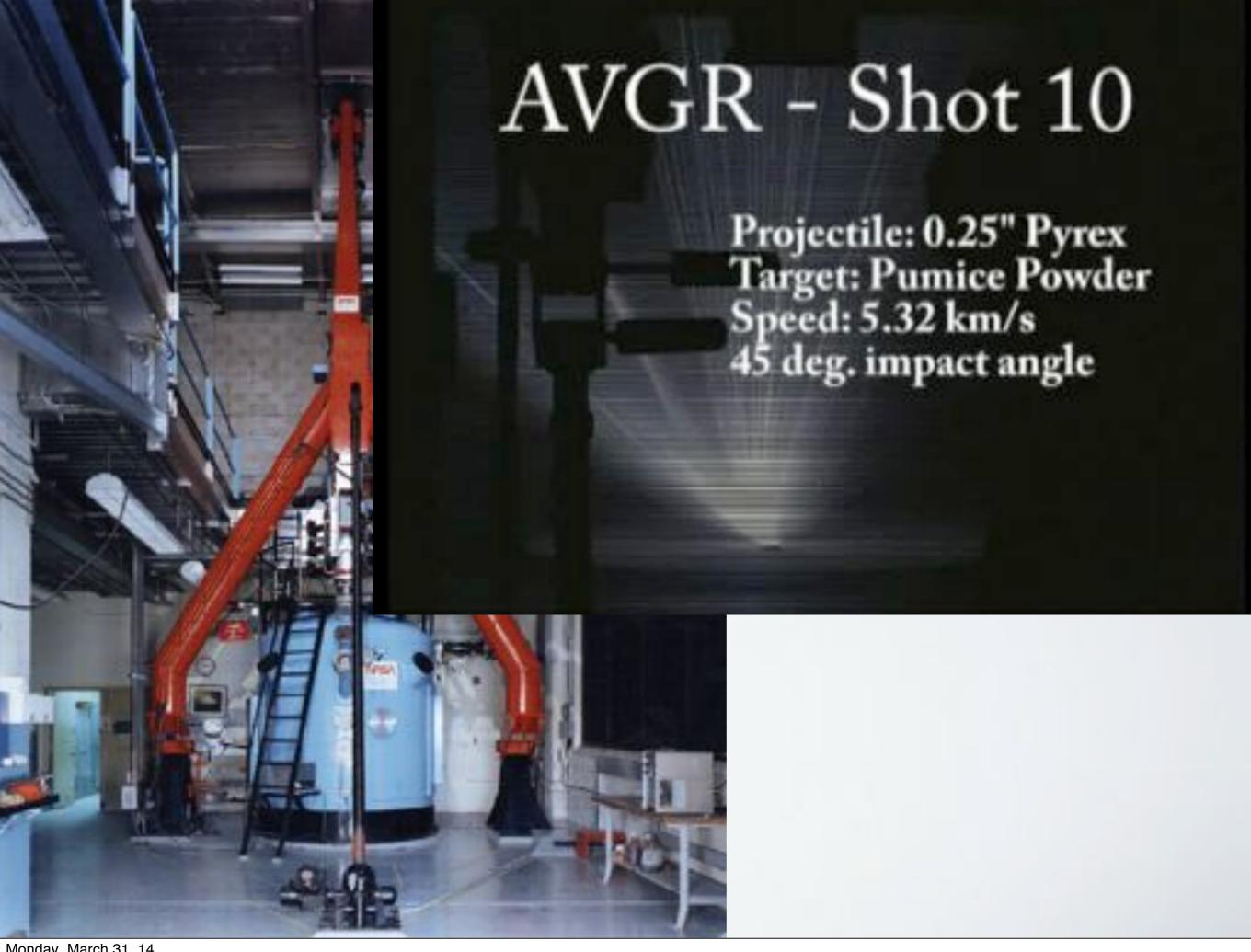
# 300+ lunar impacts observed 2005-present







Monday, March 31, 14





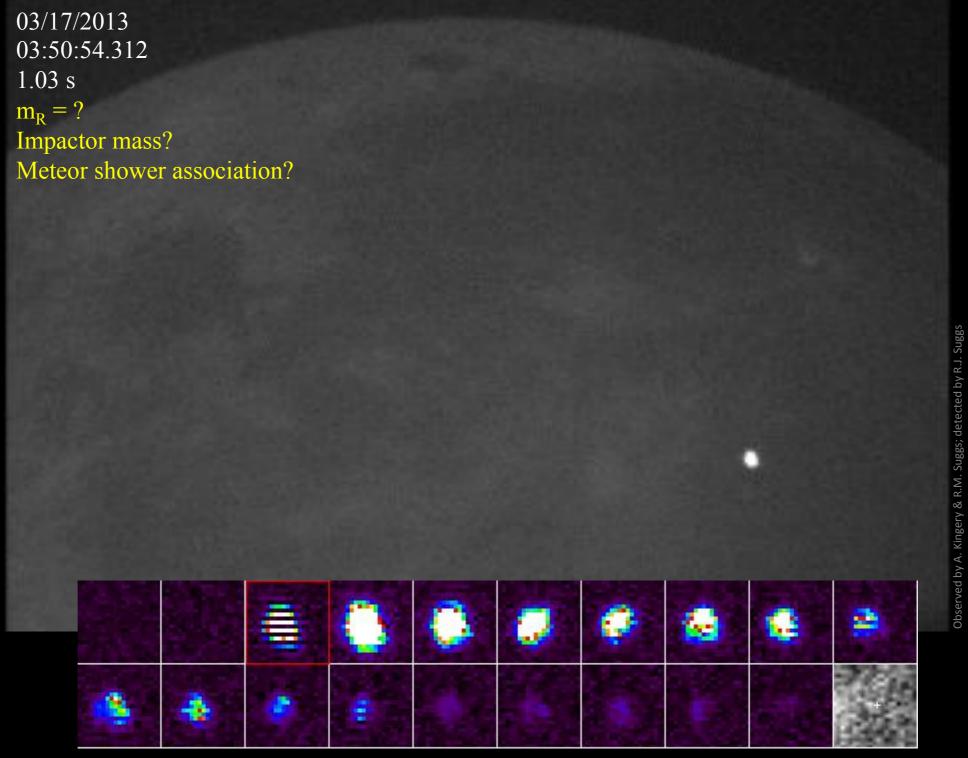
## AVGR - Shot 10

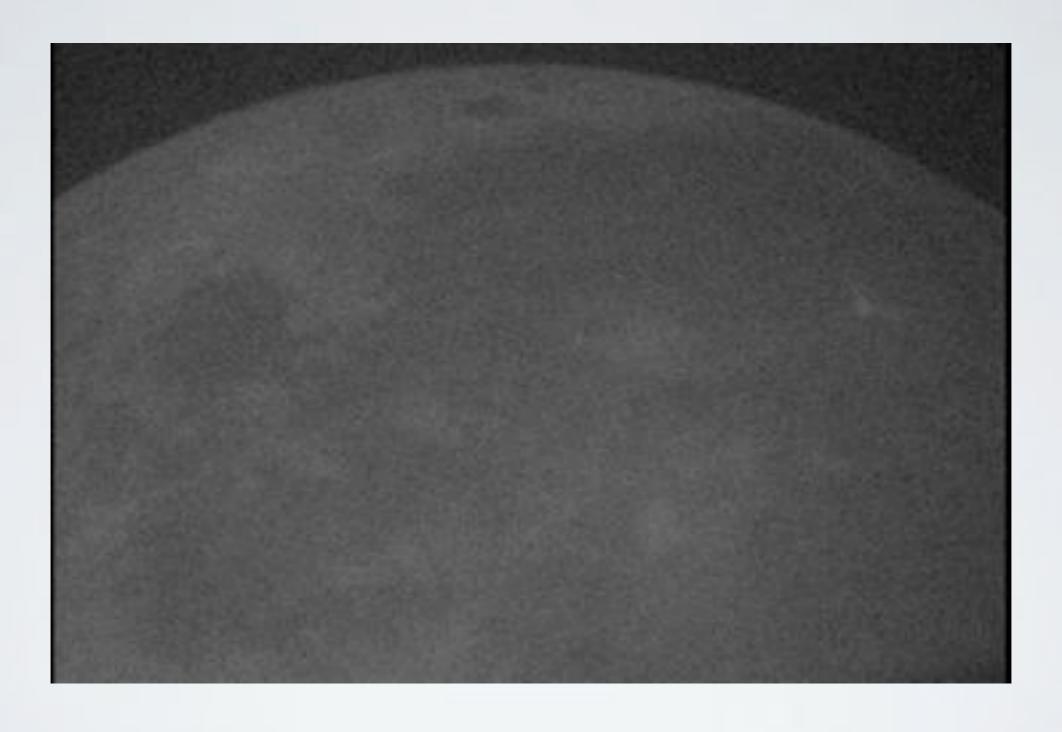
Projectile: 0.25" Pyrex Target: Pumice Powder Speed: 5.32 km/s

011 00:04:712

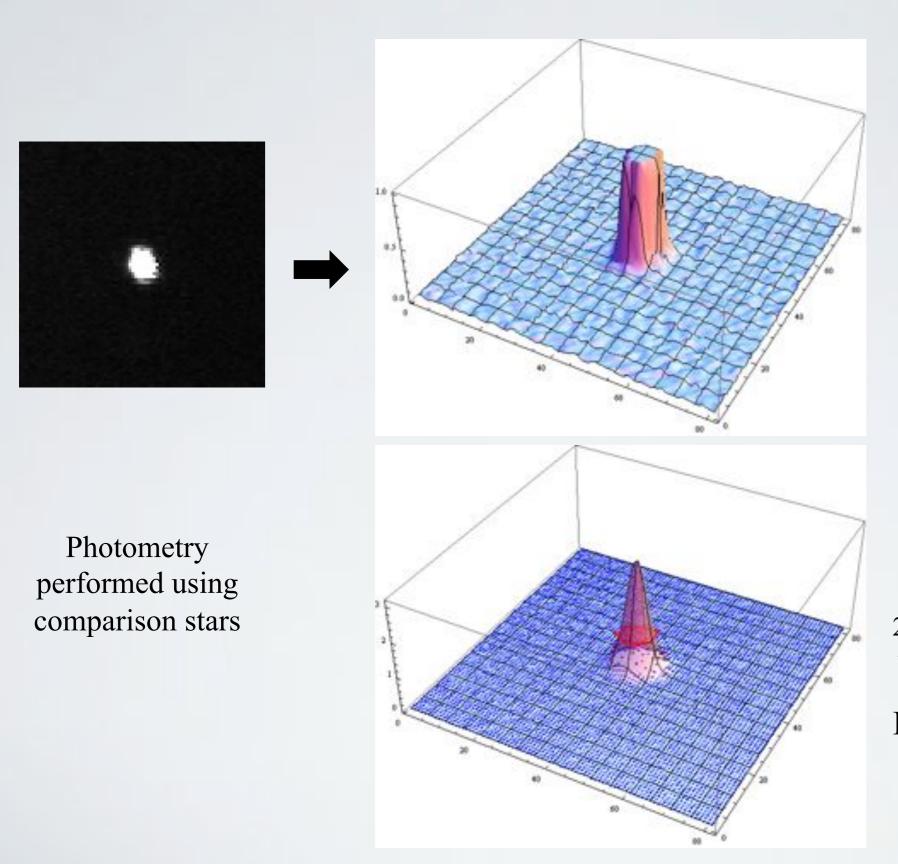
NASA AVGR

#### March 17, 2013 3:50:54 UTC





#### Estimating the magnitude



Saturated

Peak  $m_R = 4.9$ 

2D elliptical Gaussian fit to the unsaturated wings

Peak  $m_R = 3.0 \pm 0.4$ 

Similar results for 2D elliptical Moffat fit

#### Preliminary energy estimate

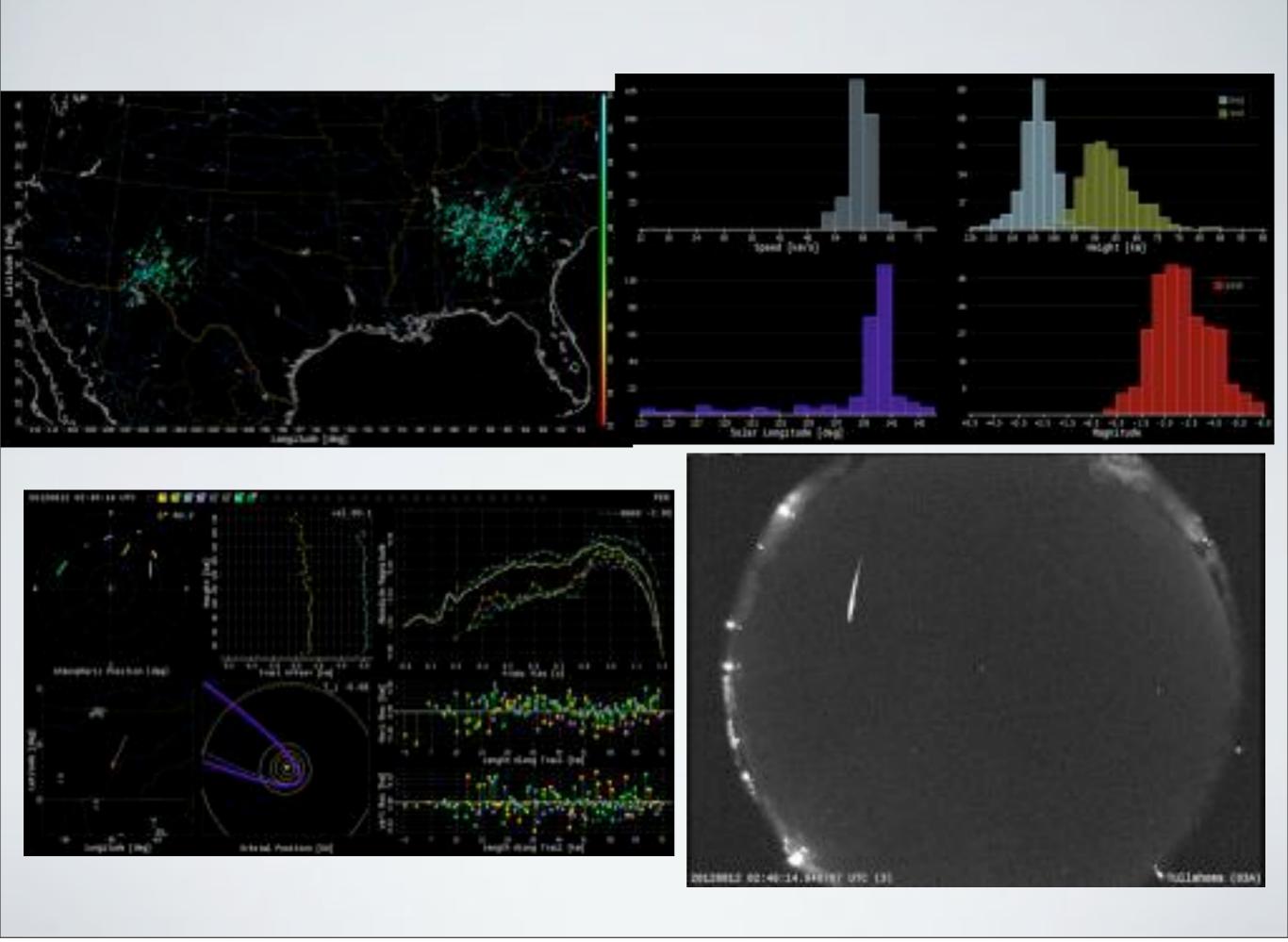
Luminous efficiency  $(\eta_{\lambda})$  relates how much of the impactor's kinetic energy (KE) is converted to luminous energy (LE) in a wavelength range  $\lambda$ 

$$LE_{\lambda} = \eta_{\lambda} KE_{\lambda}$$

	Const. $\eta = 2 \times 10^{-4}$		Vel. dep. $\eta = 1.3 \times 10^{-3}$ (Moser et al. 2011)		
	Average	Range	Average	Range	
Luminous energy (J)	$7.1 \times 10^{6}$	$4.7 \times 10^6 - 1.1 \times 10^7$	$7.1 \times 10^{6}$	$4.7 \times 10^6 - 1.1 \times 10^7$	
Kinetic energy of impactor (J)	$3.6 \times 10^{10}$	$2.4 \times 10^{10} - 5.5 \times 10^{10}$	$5.4 \times 10^{9}$	$3.6 \times 10^9 - 8.4 \times 10^9$	
Impactor mass (kg) (assuming v <sub>g</sub> = 25.6 km/s)	108	72 – 168	16	11 – 26	

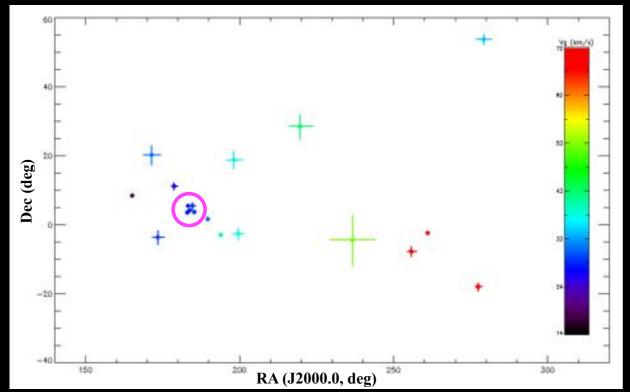
Why did we assume  $v_g = 25.6 \text{ km/s}$ ?





## Meteor shower association

19 fireballs seen on Mar 17, 2013

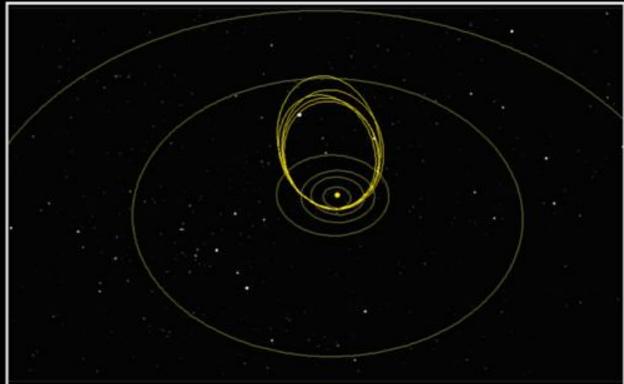


Geocentric meteor radiants color-coded by speed with a tight cluster of 5 with

Virginid Complex

	meteors	$NVI^1$	EVI <sup>2</sup>
$\alpha_{\rm g}(^{\rm o})$	$184.1 \pm 1.0$	185.7	183.6
$\delta_{\rm g}$ (°)	$4.4\pm0.9$	2.3	3.7
$v_g (km/s)$	$25.6 \pm 0.8$	23.0	28.9
$\lambda_{\text{sun}}^{\circ}$ (°)	356.6	354	354

Cluster of 5 seen on Mar 17, 2013

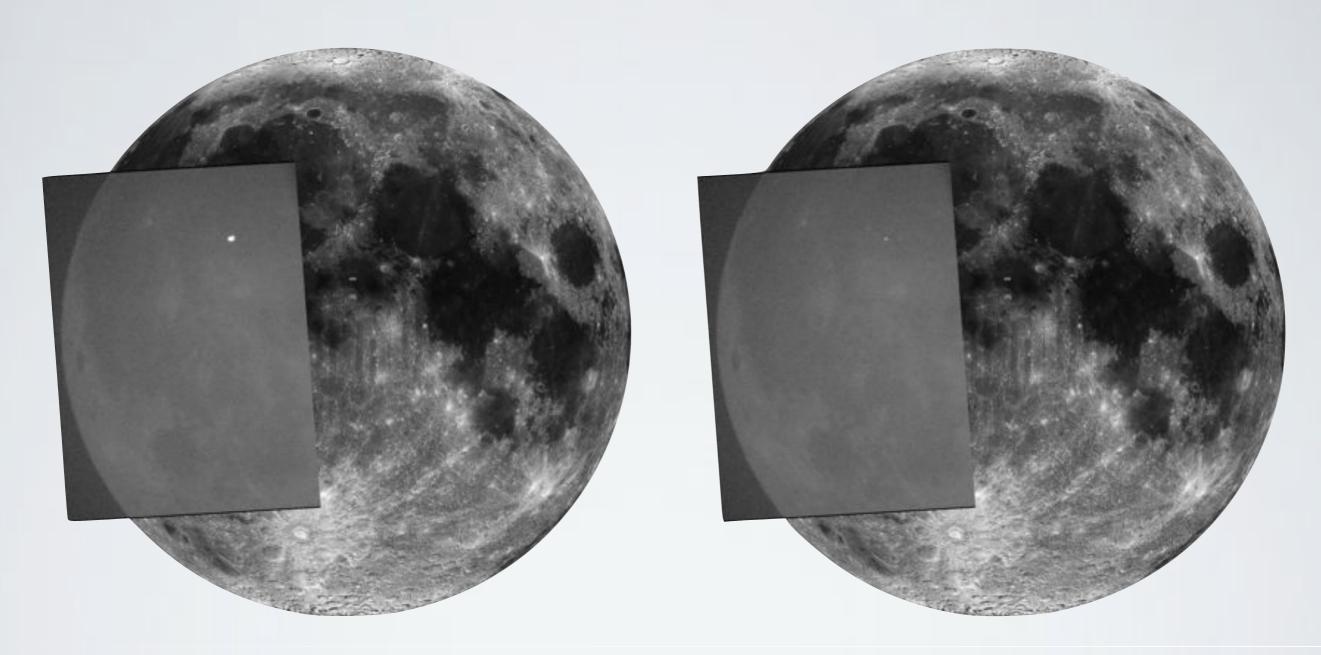


Orbits of the cluster of 5 were very similar with the following average orbital elements

	meteoroids	NVI	EVI
a (AU)	$2.25 \pm 0.17$	1.69	2.82
e	$0.79 \pm 0.02$	0.71	0.86
i (°)	$5.26 \pm 1.02$	3.7	5.2
$\omega$ (°)	$280.32 \pm 2.11$	282.4	285.8
$\Omega\left(^{\circ}\right)$	$356.65 \pm 0.07$	358.0	355.1
q (AU)	$0.48 \pm 0.02$	0.496	0.40
Q (AU)	$4.0 \pm 0.3$	2.89	5.25
Tj	$3.1 \pm 0.2$	Indicate	es asteroidal boo

<sup>1</sup>(Sekanina, 1973), <sup>2</sup>(Whipple, 1957)

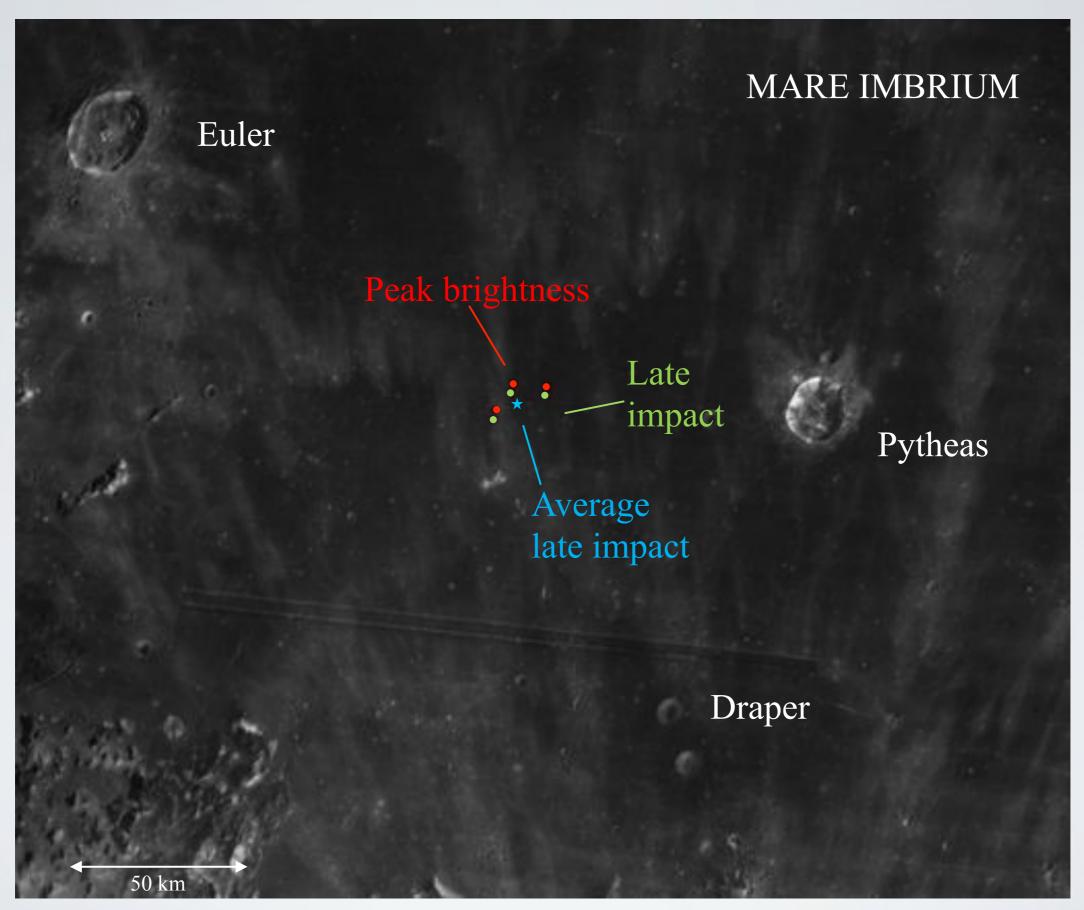
## Mapping the impact location



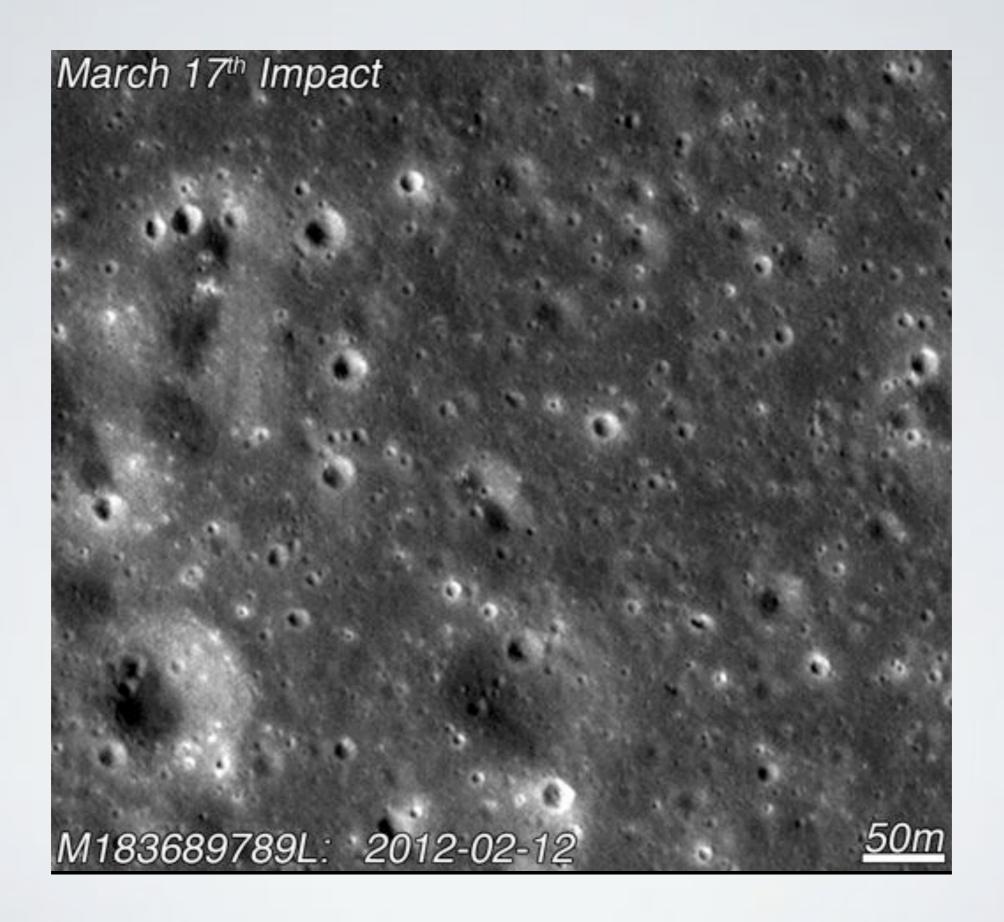
Flash at peak brightness

Flash 10 frames (333 ms) after the peak

ArcMap (ArcGIS 10) was used to georeference the lunar impact video



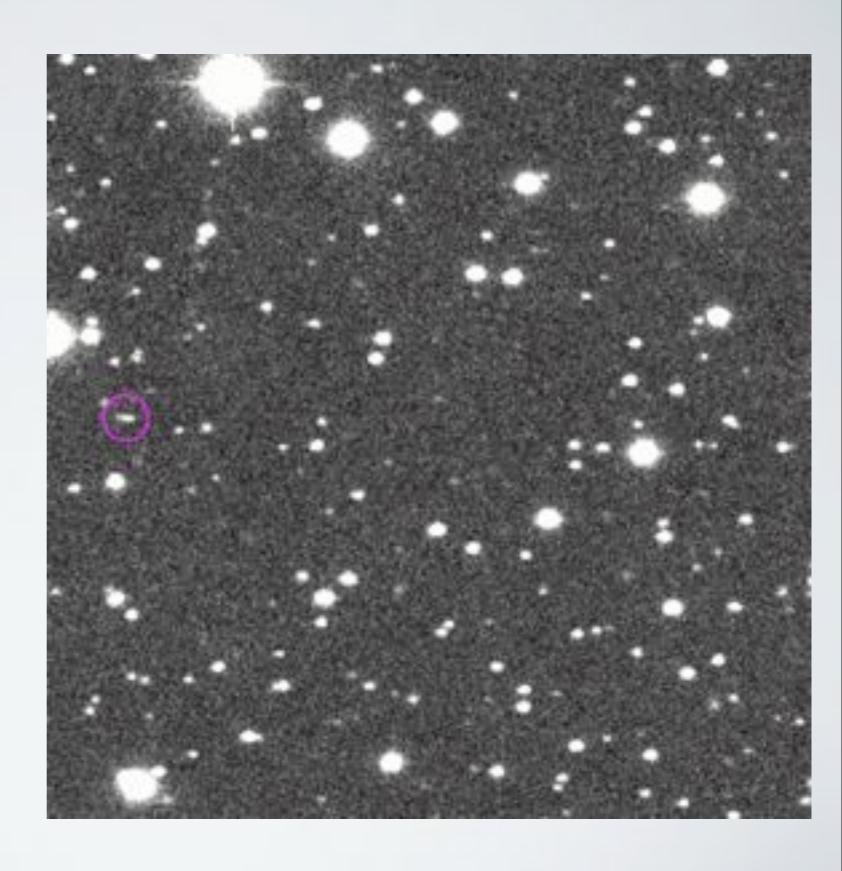
Average location:  $20.599 \pm 0.172^{\circ} \text{ N}$ ,  $23.922 \pm 0.304^{\circ} \text{ W}$ 





## 2014 AA

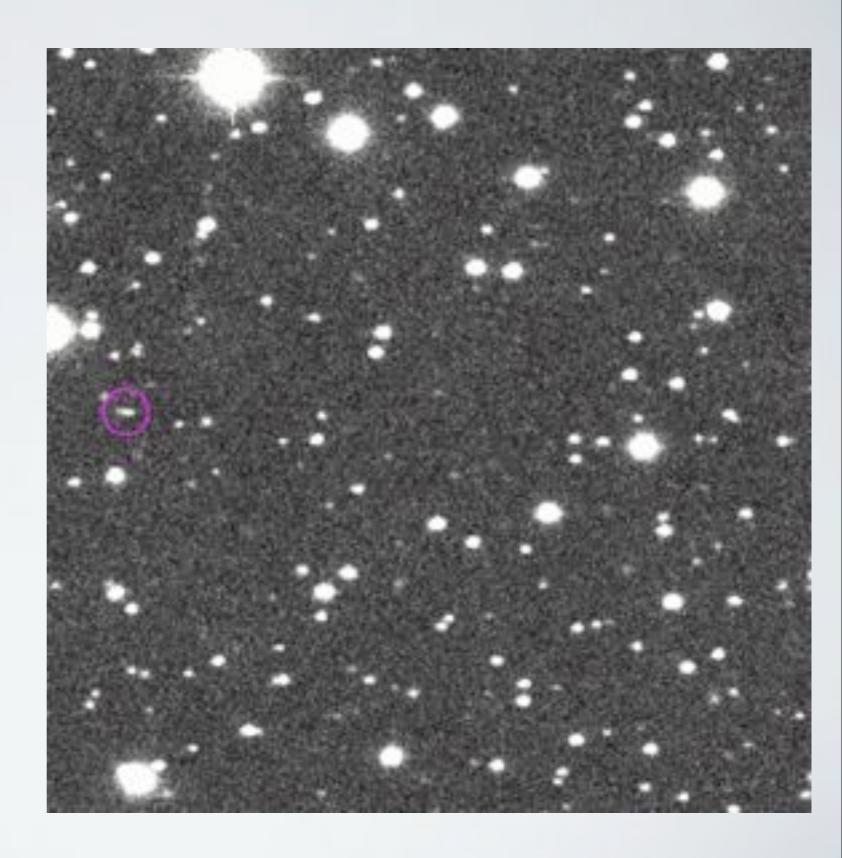
Discovered January 1 at 6:18 UT. 7 images taken over 70 minute interval



#### 2014 AA

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Magnitude of 19 indicated an object 2-4 meters in diameter

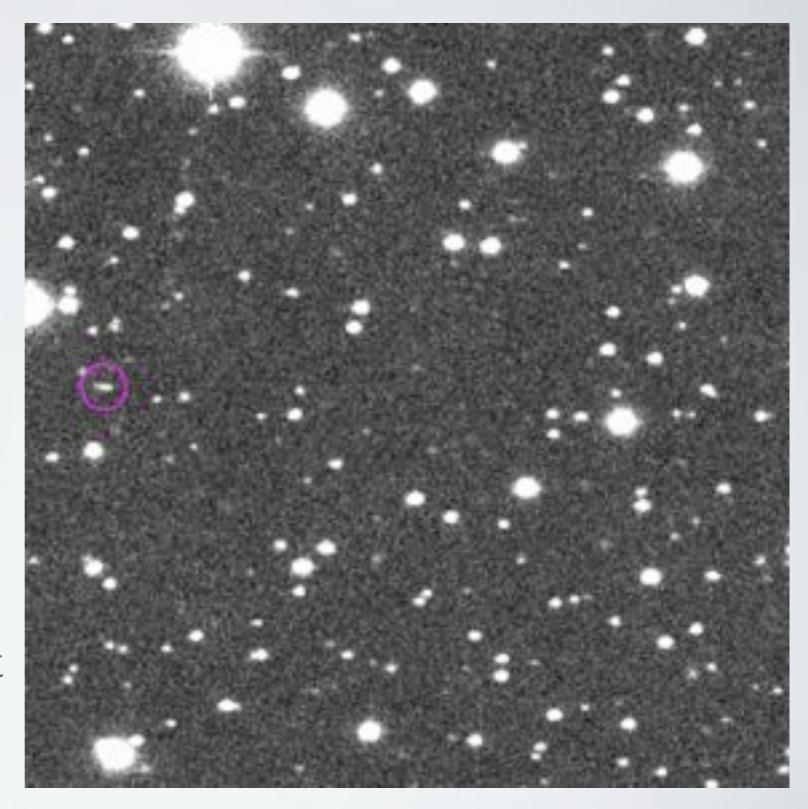


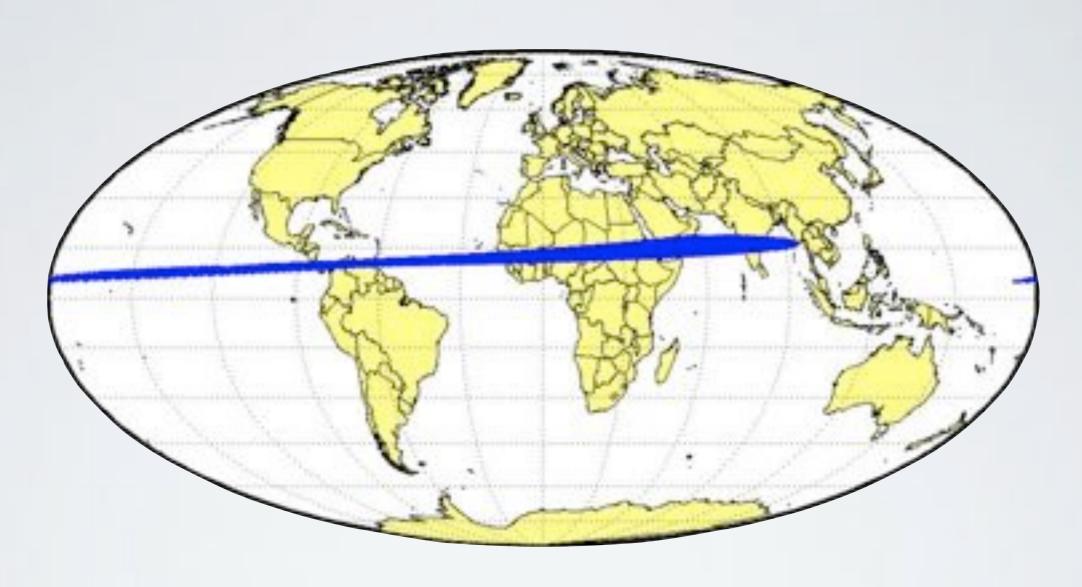
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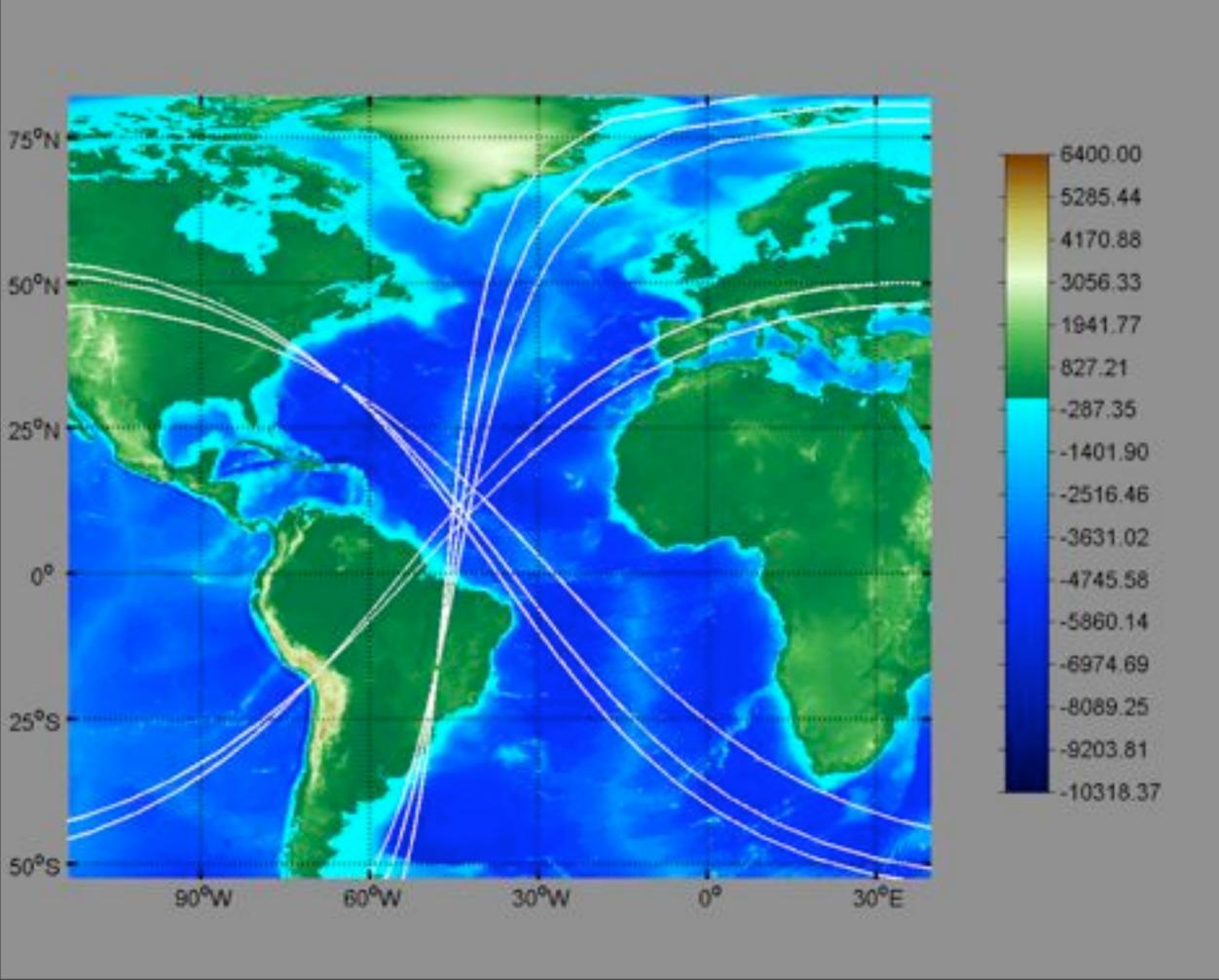
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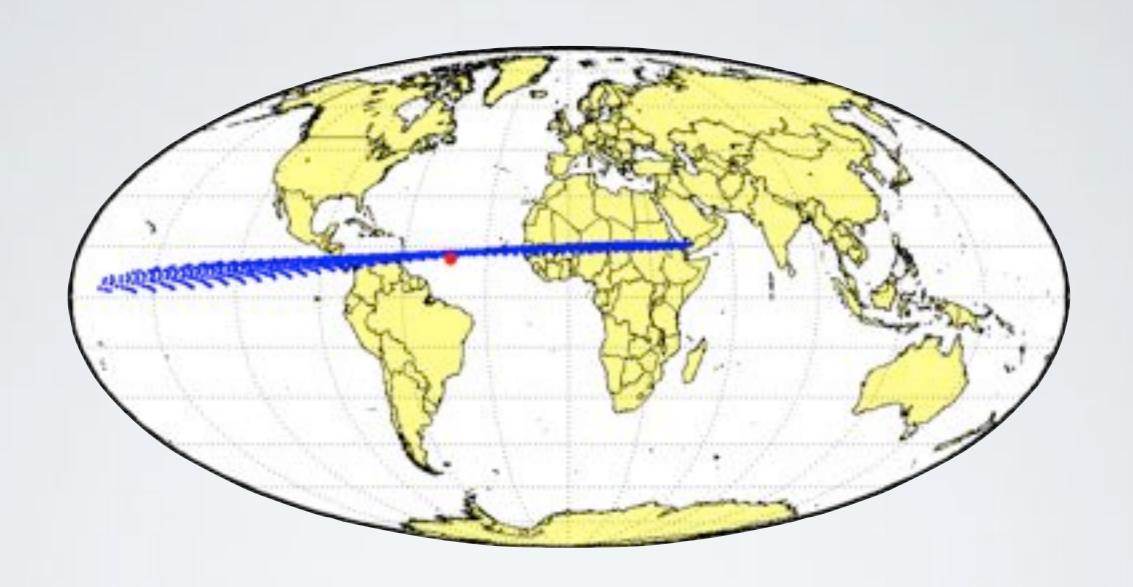
Positions submitted to the MPC, but no one realized object was on an impact course with Earth until just before impact



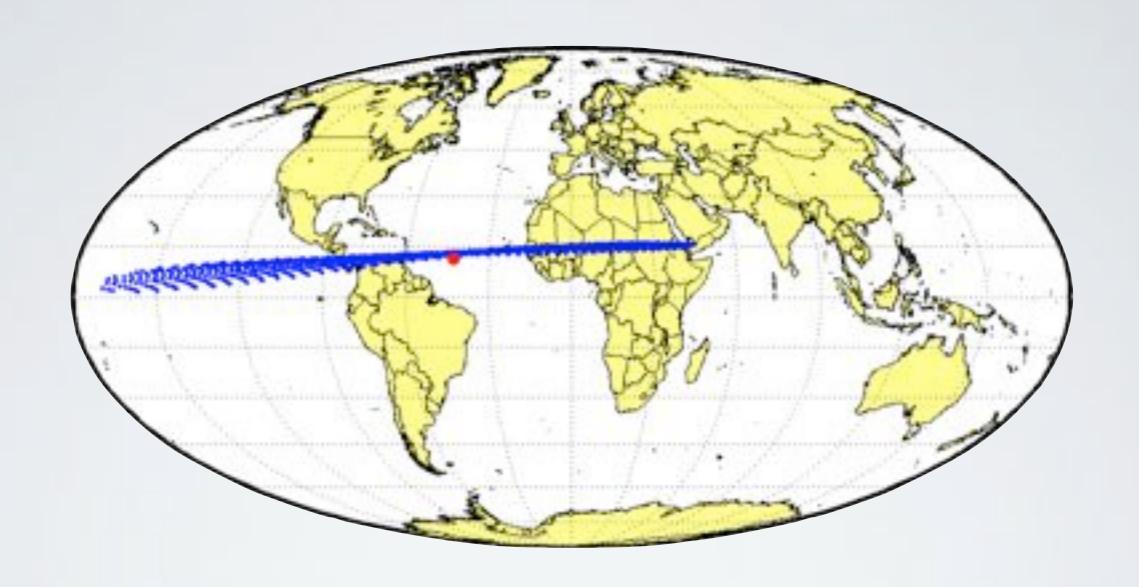


Projected impact time was Jan 2 2:25 UTC, just off coast of west Africa





Actual impact time was Jan 2 3:30 UTC +/- 30 minutes



Actual impact time was Jan 2 3:30 UTC +/- 30 minutes Impact energy of about 1 kiloton

# CHELYABINSK

## Fireball Properties

Energy 470 kilotons at altitude of 30 km (19 miles)

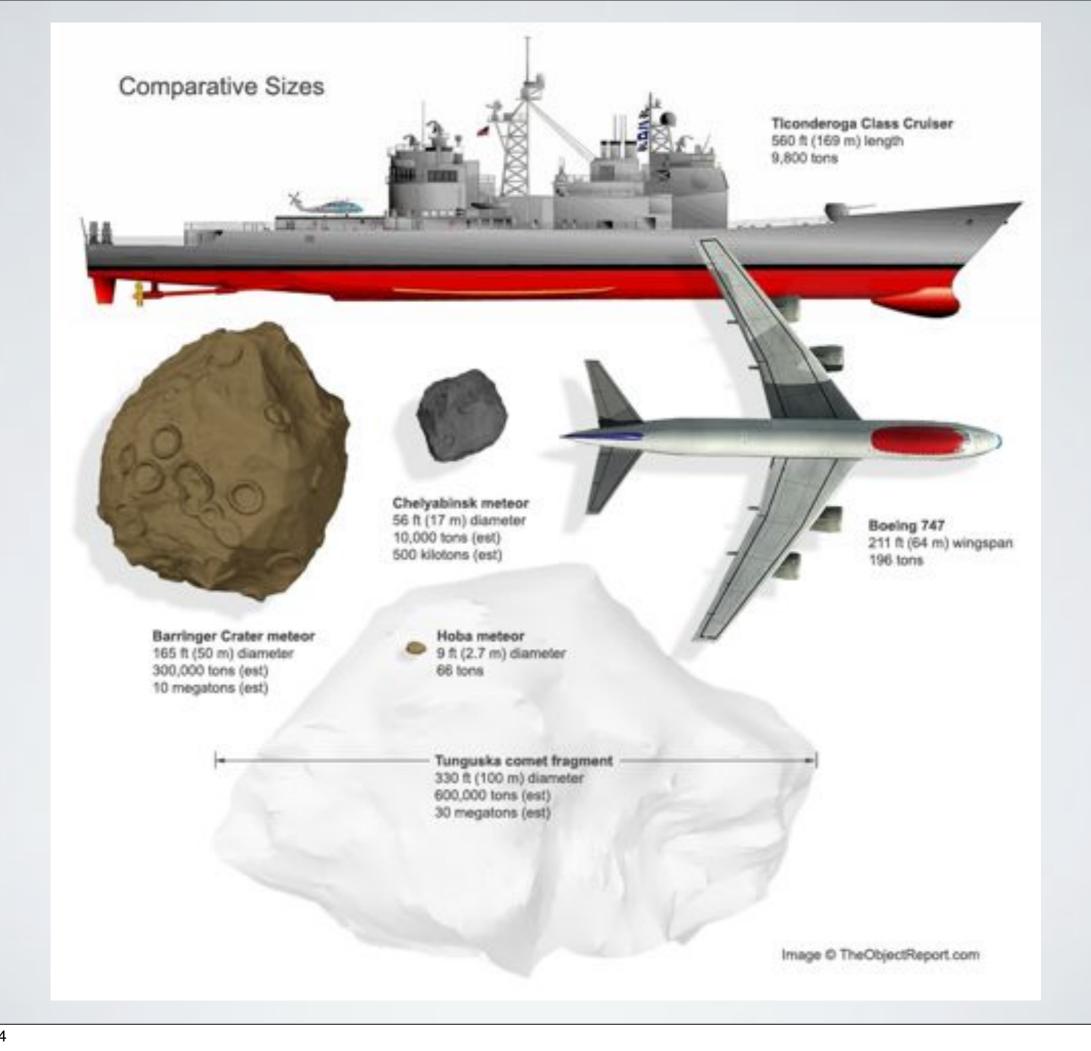
near Chelyabinsk (54.8° N, 61.1° E)

**Speed** 19 km/s (42,500 mph)

Mass/size ~12,000 tons

19 m (62 ft)

**Composition** Ordinary chondrite (LL<sub>5</sub>)









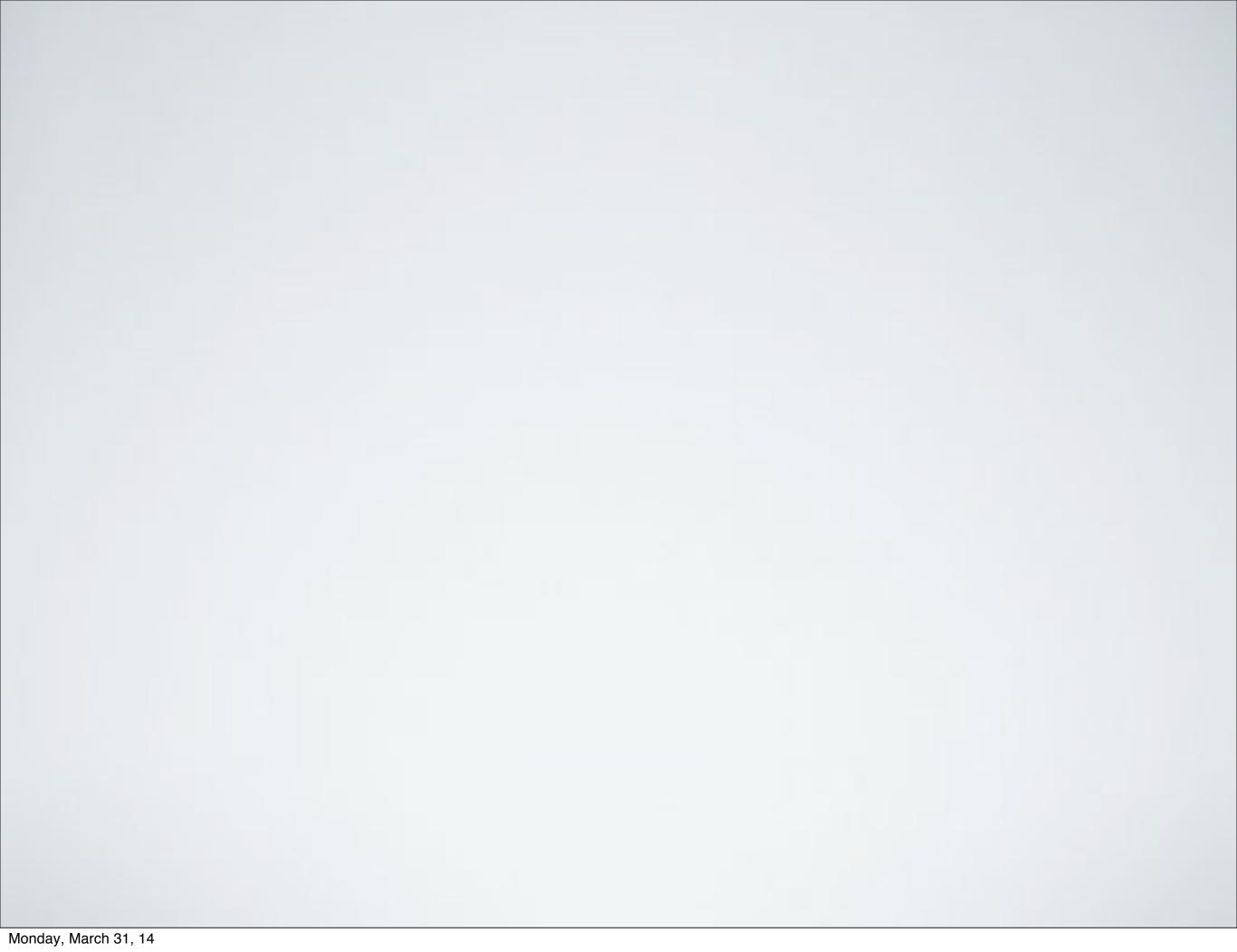


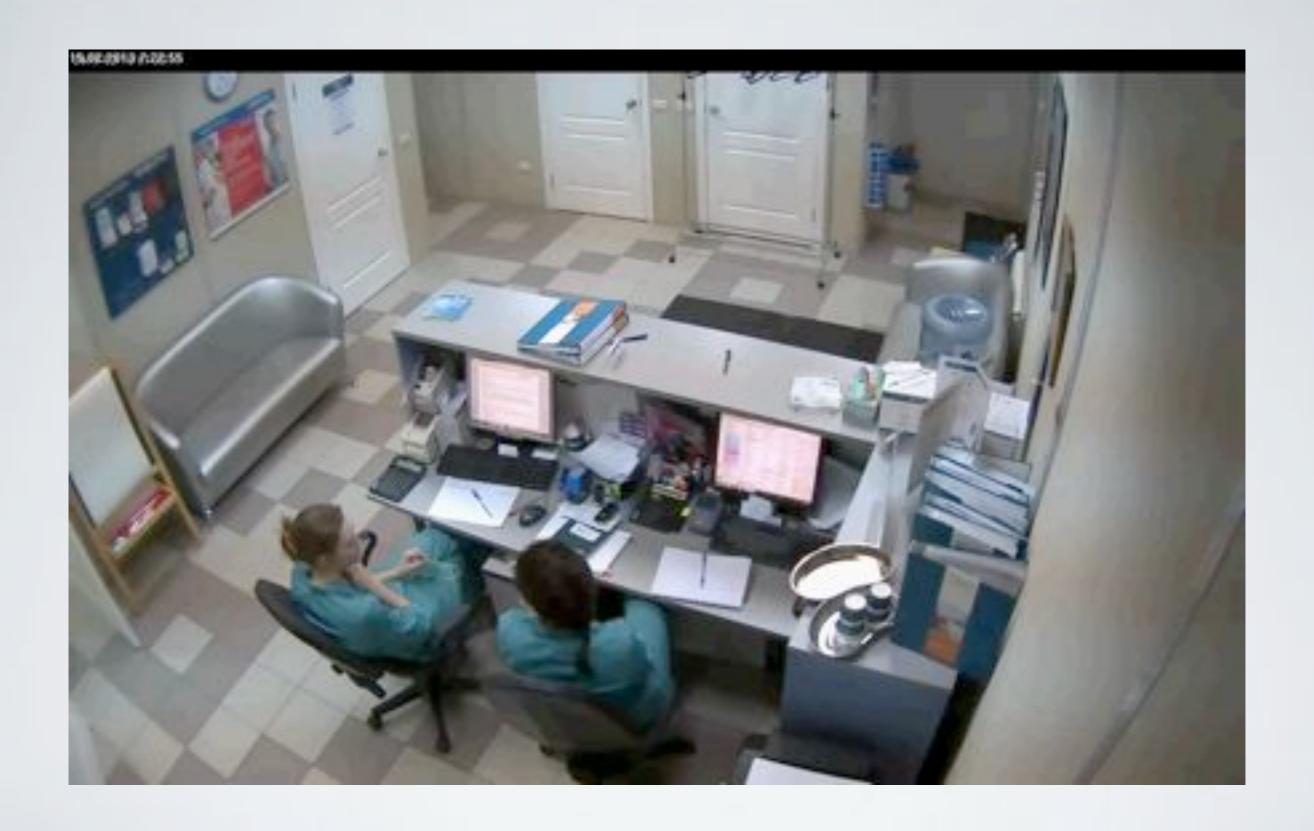


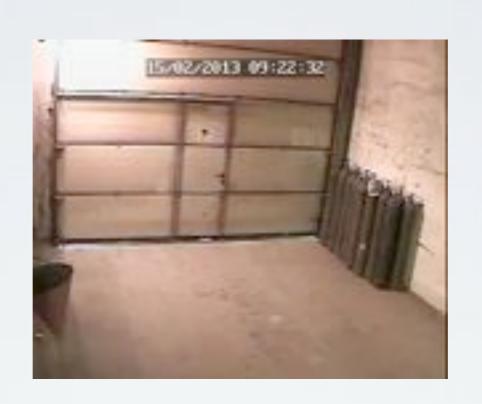












#### Kiloton Airbursts Detected by Infrasound: 2000-2013

